

# COMING-TAO Collaboration

## “COMING-PLUS”

KANEKO Hiroyuki(NAOJ/Nobeyama Radio Observatory)  
& COMING-PLUS team

# Nobeyama Radio Observatory

- One of branches of National Astronomical Observatory of Japan
- Operates the 45-m radio telescope
- One of the largest radio telescopes in mm reign
- 20GHz(15mm)-115GHz(2.7mm)



# What's COMING?

- COMING: CO Multiline Imaging of Nearby Galaxies
- One of Legacy Projects of the observatory

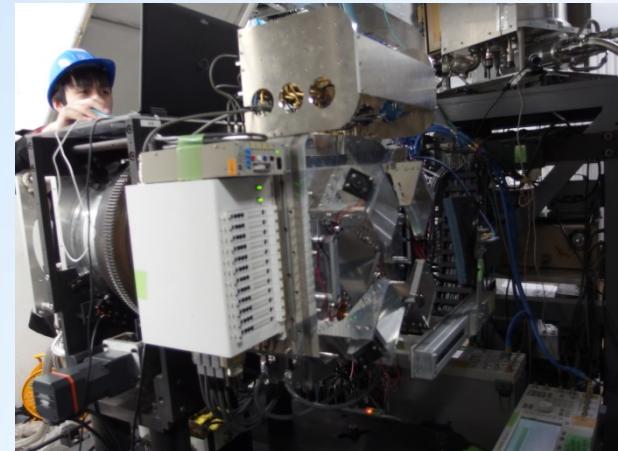
## COMING

CO MULTI-LINE IMAGING OF NEARBY GALAXIES

- Targets: Nearby Galaxies (P.I.: Sorai [Hokkaido Univ.]
- Simultaneous observations of molecular lines:  
 $^{12}\text{CO}(J=1-0)$ ,  $^{13}\text{CO}(J=1-0)$  and  $\text{C}^{18}\text{O}(J=1-0)$

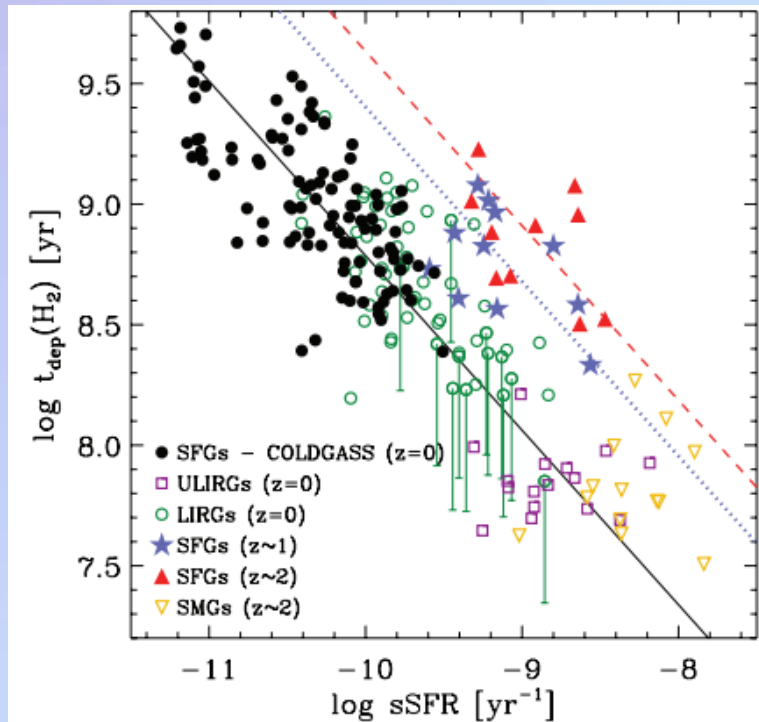
with new receiver FOREST

## FOREST



# What's COMING?

- CO molecular lines: basic tracers of molecular gas
- Fuel of star formation
- ⇒ fundamental process of evolution of galaxies
- Star formation mechanism (where and how?)

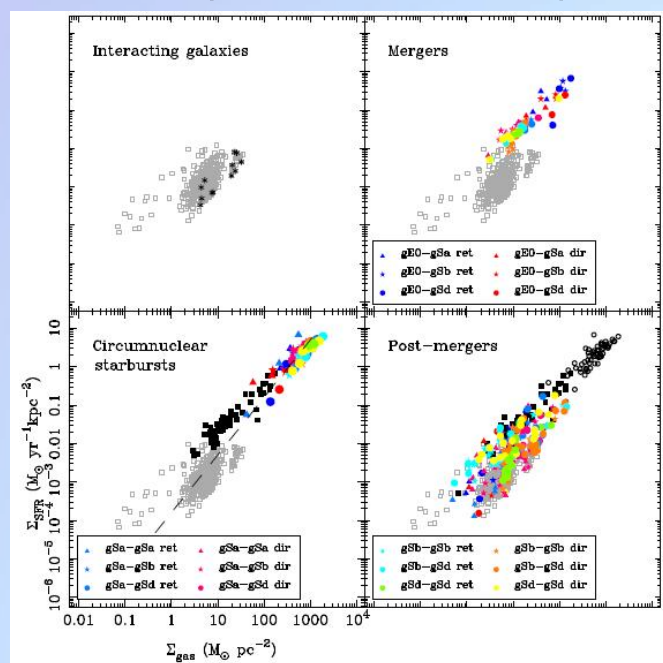


Saintonge et al. (2011)



# What's COMING?

- CO molecular lines: basic tracers of molecular gas
- Fuel of star formation
- ⇒ fundamental process of evolution of galaxies
- Relationship btw gas properties and environments (bar, AGN, starburst, interactions, cluster, etc.)



Matteo et al. (2008)



# What's COMING?

- Comparison to previous CO imaging surveys
  - BIMA SONG: 44 galaxies @ 6"
  - Nobeyama CO atlas: 40 galaxies @ 16"
  - STING: 23 IR-bright galaxies @ 3-5"
- ⇒ few # of samples, low spatial resolution



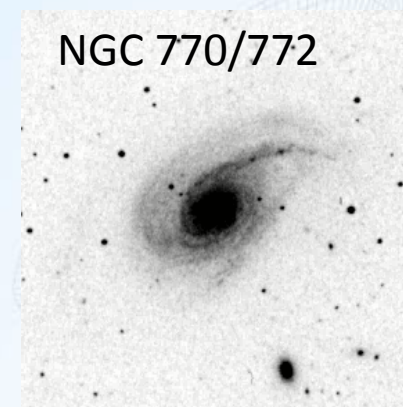
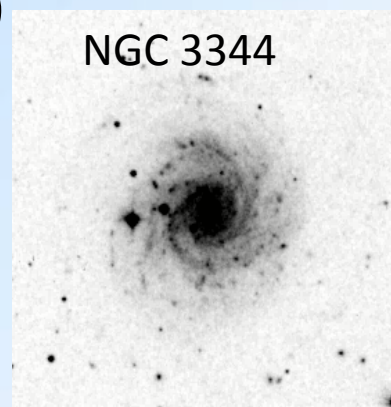
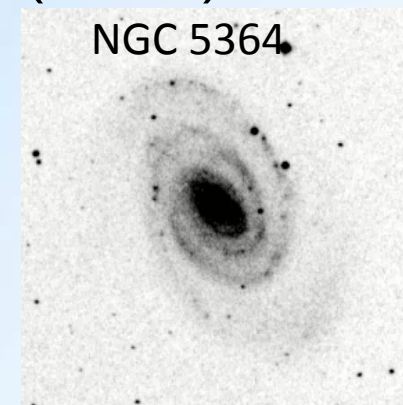
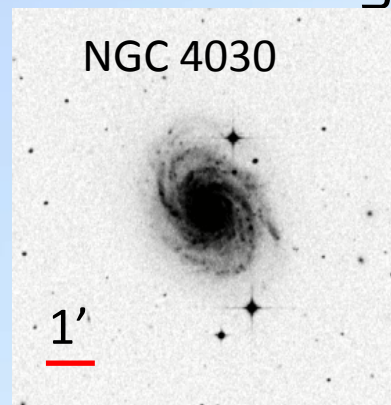
# What's COMING?

- galaxies with the distance of  $< 35$  Mpc
- # of targets: 238 galaxies
- ...more than the # of galaxies ever imaged ( $\sim 100$ )

- Typical galaxy size :  $5'$
- Imaging area: 70% of  $D_{25}$
- Velocity res.: 10 km/s
- Spatial res.:  $19''$  @  $^{12}\text{CO}(1-0)$

- $^{12}\text{CO}(J=1-0)$ ,  $^{13}\text{CO}(J=1-0)$   
and  $\text{C}^{18}\text{O}(J=1-0)$

→ gas density and temperature



# What's COMING?

- From low- $z$  to high- $z$
- If a target at  $z \sim 1$  is observed w/ ALMA's highest res.  
...linear res.  $\sim$  kpc  
 $\Rightarrow$  comparable to the observation of nearby galaxies  
( $\sim$  few 10 Mpc) w/ the 45-m telescope
- Enable us to direct comparison
- Important to understand the statistical properties of low- $z$  galaxies with resolving their structures



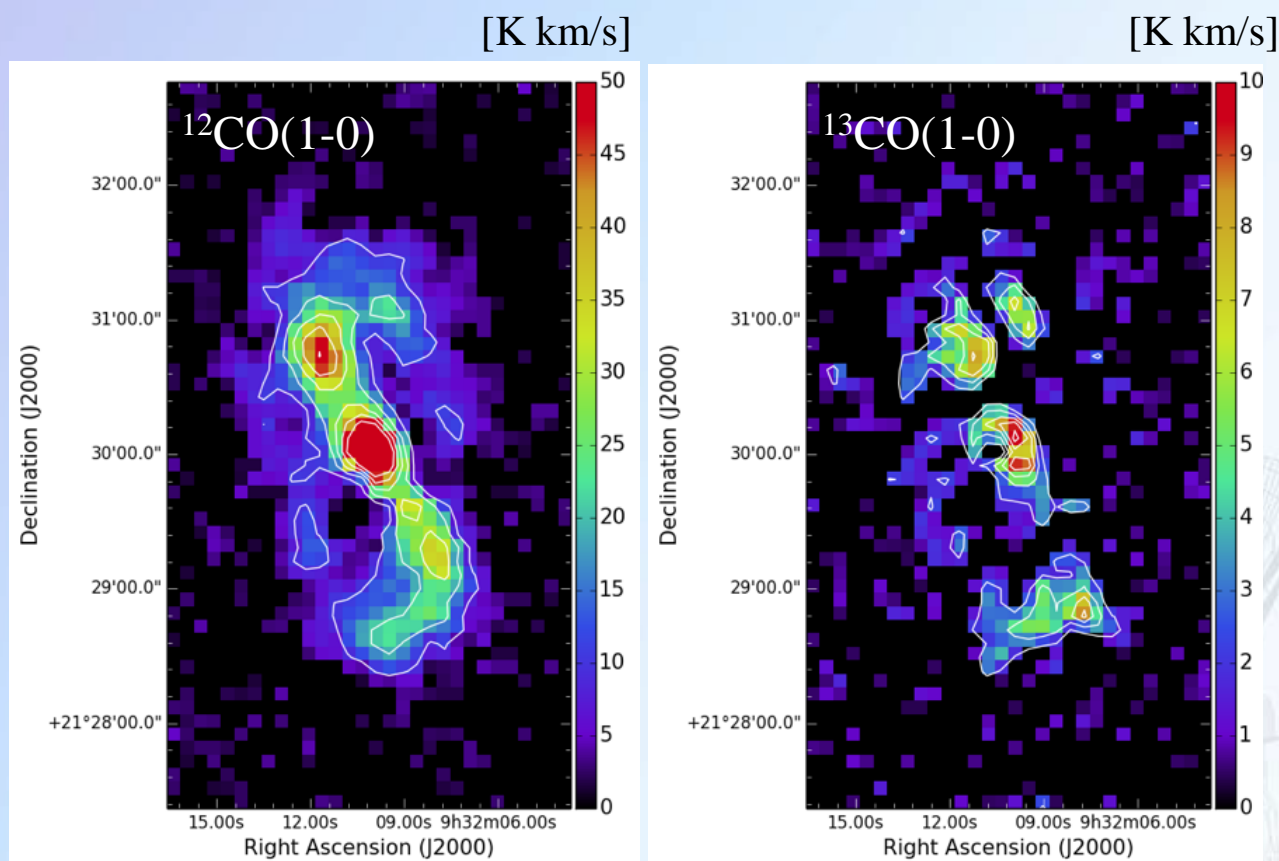


# "COMING" has come!

- From this spring, scientific data has obtained
- Successfully obtained  $^{12}\text{CO}(1-0)$  &  $^{13}\text{CO}(1-0)$  maps simultaneously!!

NGC 2903

Optical

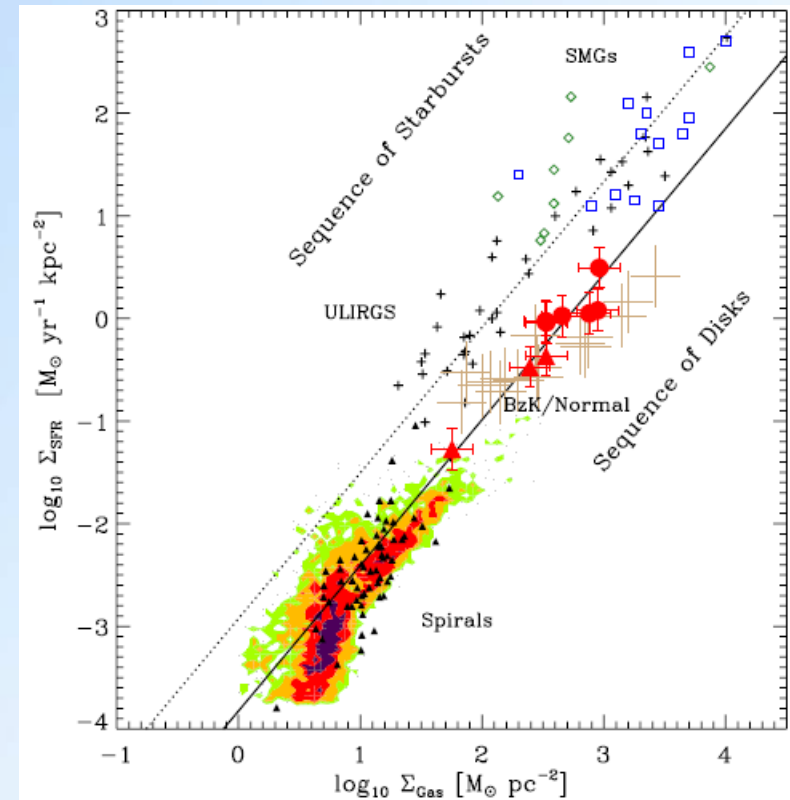


# COMING-PLUS

- Relationship btw molecular gas and star formation
- Kennicutt-Schmidt law (e.g., Kennicutt, 1998)
- One of the important topics in the studies of evolution of galaxies

$$\Sigma_{\text{SFR}} \propto \Sigma_{\text{gas}}^N$$

- $N \sim 1-2$
- $N$  differs in galaxies and structures (e.g., Daddi et al. 2010)



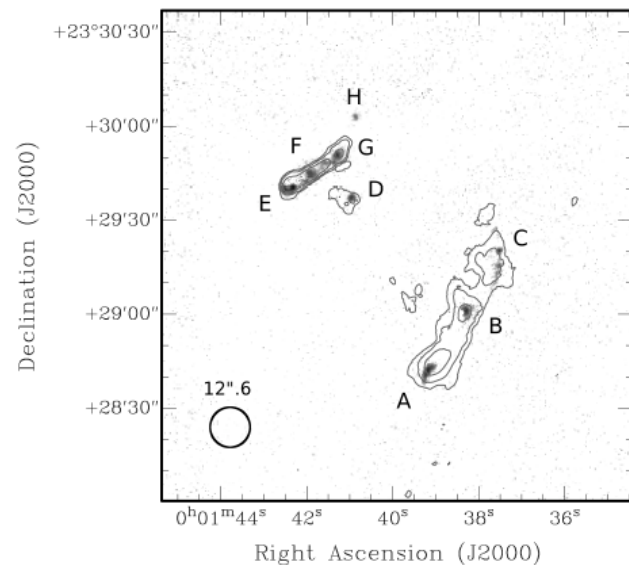
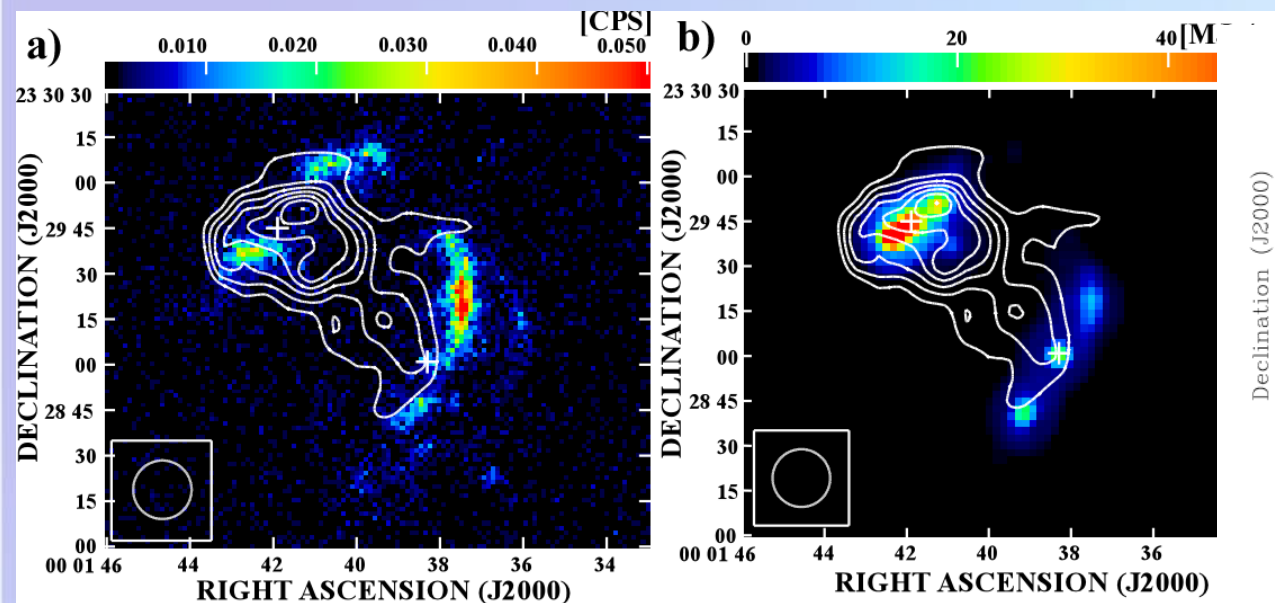
# COMING-PLUS

- Example of CO and Pa $\alpha$ 
    - …for science w/ imaging data, H $\alpha$  is mostly used
  - Problem: dust extinction
    - correction w/ dust re-radiation (e.g., Calzetti et al. 2008)
      - Worse in spatial res. (3-4" )
      - Needs of two data set (H $\alpha$  & dust [24  $\mu$ m, etc.]
  - When science w/ ALMA, resolution is NOT enough
    - …Pa $\alpha$  should be a clue.
- miniTAO/ANIR and TAO/SWIMS are the best answer**



# COMING-PLUS

- Tracers of star formation
- Interacting galaxies VV 254



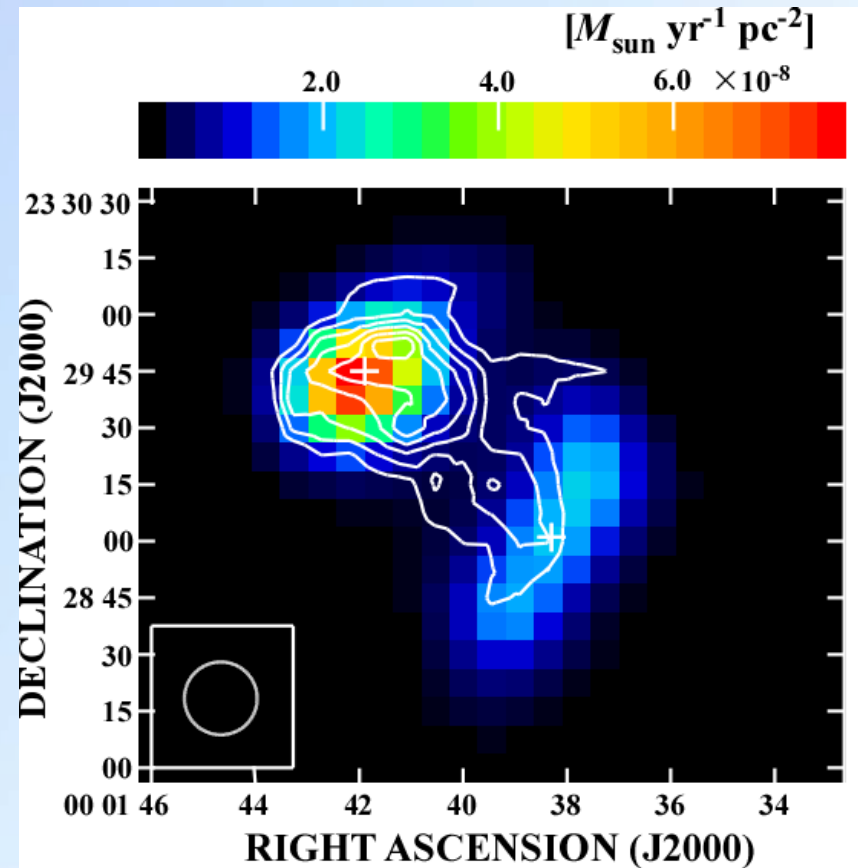
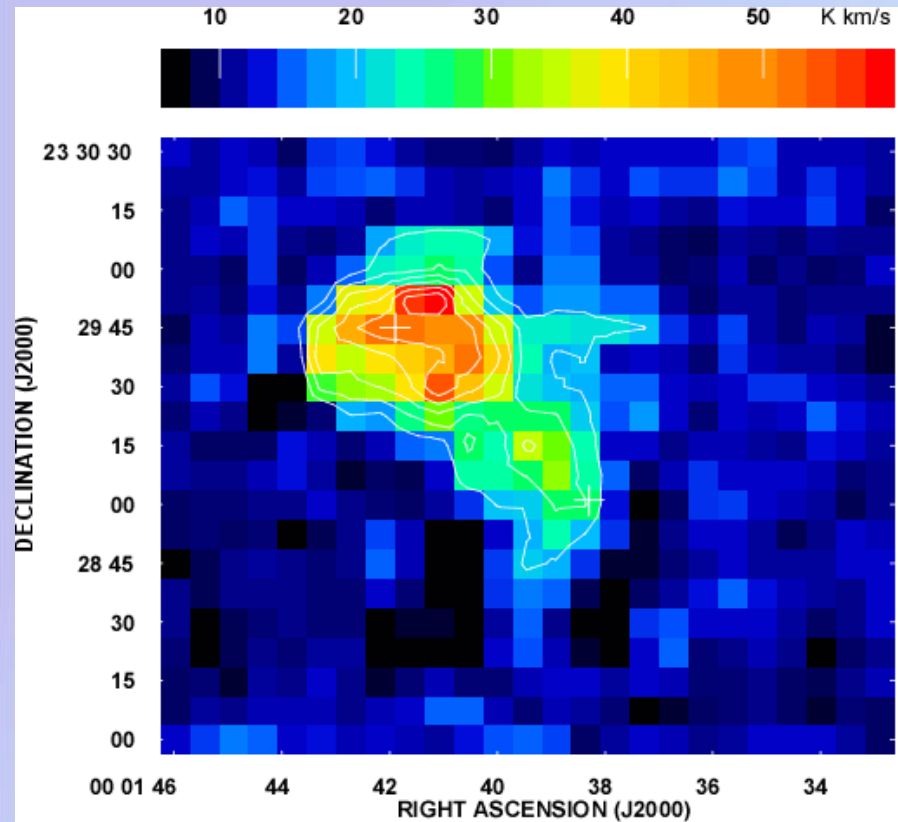
FUV(GALEX)

Spitzer 24 um

Komugi et al. (2012)

# COMING-PLUS

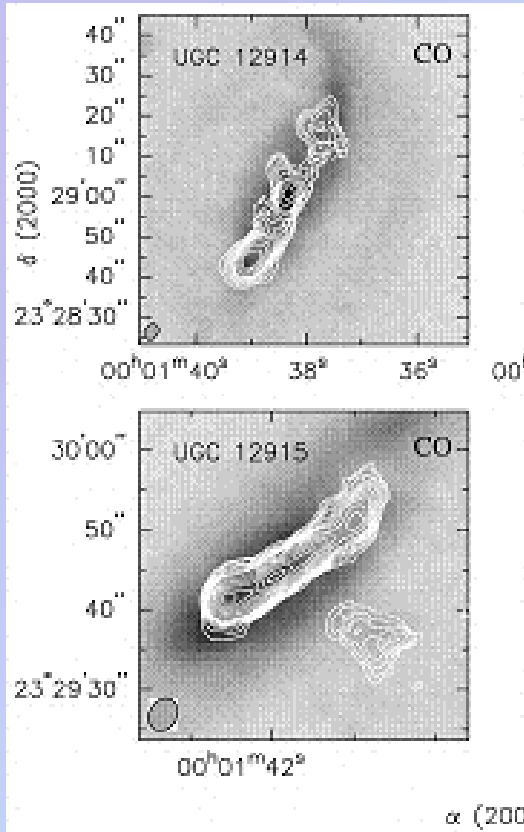
- Tracers of star formation
- Interacting galaxies VV 254



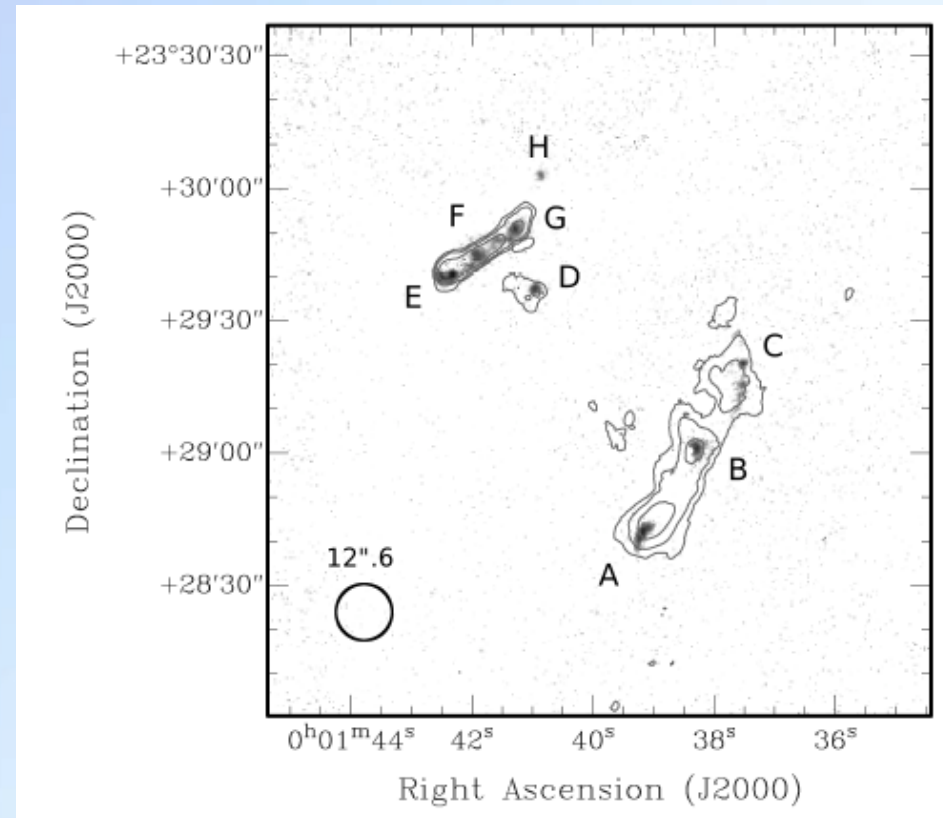
$^{12}\text{CO}$  (45-m: Kaneko et al. 2013) SFR from FUV + 24um (Kaneko et al. in prep.)

# COMING-PLUS

- Tracers of star formation
- Interacting galaxies VV 254



$^{12}\text{CO}$  (OVRO: Iono et al. 2005)



$\text{Pa}\alpha$  (Komugi et al. 2012)

contour :  $^{12}\text{CO}$ (Iono et al. 2005 as left fig.)

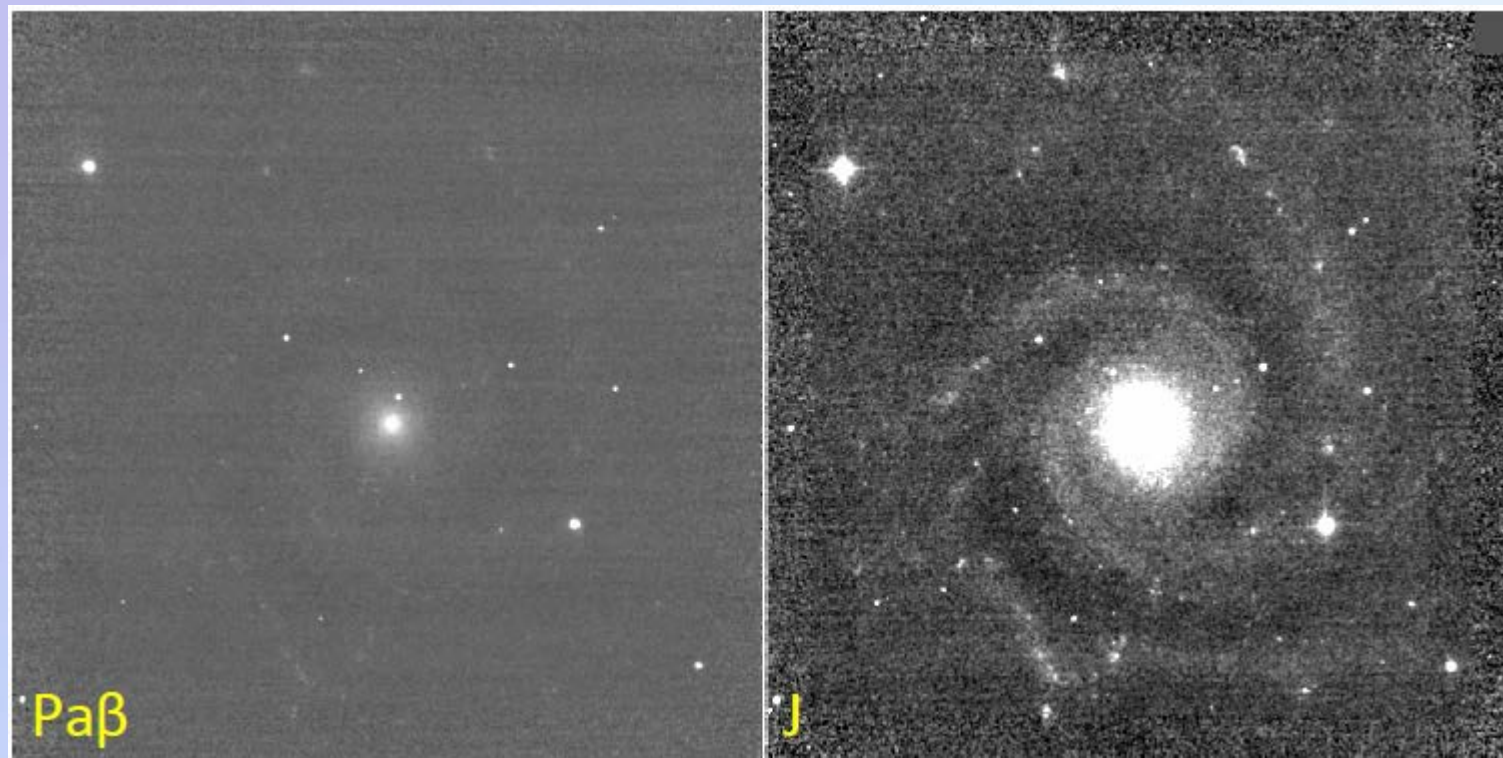
# COMING-PLUS

- COMING containing Pa $\alpha$  Line imaging by Using the telescopes in Southern hemi-sphere
- Pa $\alpha$  imaging observations
- Targets: galaxies in the sample of COMING (238 galaxies) which can be observable from the TAO site (dec. < 30 degree)  
→ candidates : 140 galaxies
- Observing time per one galaxy (rough estimate FOR miniTAO)  
broad band (J , H + Ks): ~10 min x 2 band  
Pa $\alpha$  , Pa $\beta$  ~40 min  
... ~1 hour/galaxy  
→ ~140 hours



# COMING-PLUS

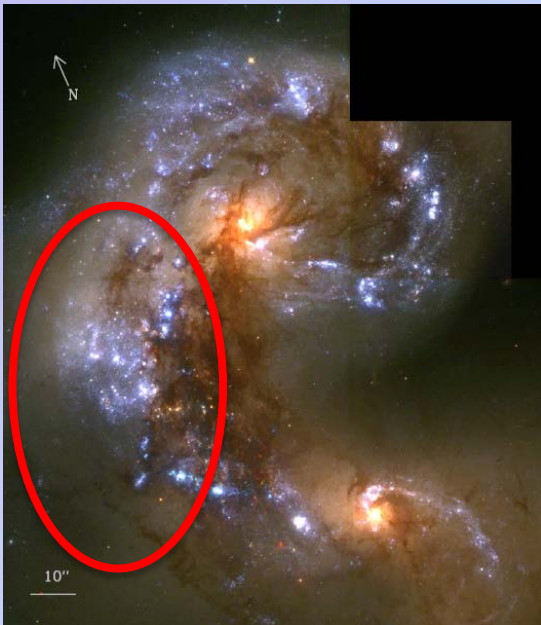
- Few galaxies: already observed w/ miniTAO/ANIR
- NGC 628



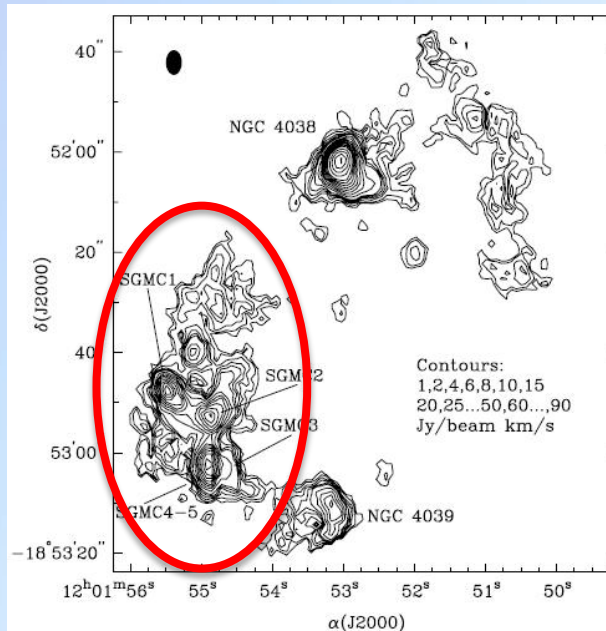


# COMING-PLUS

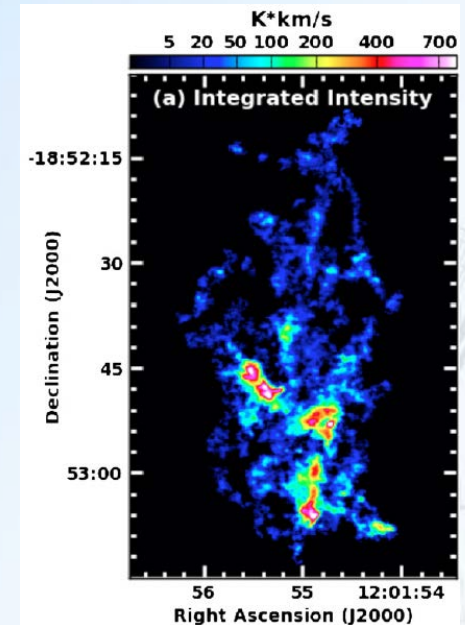
- Strength of COMING-PLUS
- molecular gas and star formation tracer images
- Easily leads to ALMA science
  - Giant Molecular Clouds scale ( $\sim 100$  pc) in nearby galaxies
  - kpc scale in High- $z$



Whitmore et al. (1999)  
 18<sup>th</sup> September, 2015



Wilson et al. (2000)  
 TAO/SWIMS Science Workshop



Whitmore et al. (2014)

# Summary

- CO multi-line imaging of nearby galaxies using the 45-m telescope, COMING, has been conducted
- Kennicutt-Schmidt law is one of clues to reveal evolution of galaxies
- Pa $\alpha$  is the best star formation tracer
- miniTAO/ANIR Pa $\alpha$  survey as a first step has started
- TAO/SWIMS & COMING collaboration will be one key to investigate the evolution of galaxies
- This collaboration leads to the science with ALMA in both local and high- $z$  universe

